

# $\epsilon_b$ Constraints on the Minimal $SU(5)$ and $SU(5) \times U(1)$ Supergravity Models

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## Abstract

We have performed a systematic analysis to compute the one-loop electroweak corrections to the  $Z \rightarrow b\bar{b}$  vertex in terms of  $\epsilon_b$  and  $R_b$  in the context of the minimal  $SU(5)$  and no-scale  $SU(5) \times U(1)$  supergravity models. With the measured top mass,  $m_t = 174 \pm 10_{-12}^{+13}$  GeV, recently announced by CDF, we use the latest LEP data on  $\epsilon_b$  and  $R_b$  ( $\equiv \frac{\Gamma(Z \rightarrow b\bar{b})}{\Gamma(Z \rightarrow \text{hadrons})}$ ) in order to constrain further the two models. We find that the present LEP data on  $\epsilon_b$  and  $R_b$  constrain the two models rather severely. Especially, the low- $\tan\beta$  region is constrained more severely.  $\tan\beta \lesssim 2.5$  (4.0) is excluded by  $\epsilon_b$  at 90% C. L. for  $m_t \gtrsim 170$  (180) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) supergravity. Even more stringent constraint comes from  $R_b$ . It excludes  $\tan\beta \lesssim 4.0$  at 90% C. L. for  $m_t \gtrsim 160$  (170) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) supergravity. We also find that the sign on  $\mu$  in the

two models can be determined from  $\epsilon_b$  and  $R_b$  at 90% C. L.

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## I. INTRODUCTION

With the increasing accuracy of the LEP measurements, it has become more important than ever performing the precision test of the standard model (SM) and its extensions. A standard model fit to the latest LEP data yields the top mass,  $m_t = 178 \pm 11_{-19}^{+18}$  GeV [1]. With this large top quark mass, the  $Z \rightarrow b\bar{b}$  vertex contribution, which is proportional to  $m_t^2$ , becomes more significant, and can provide a powerful tool to constrain  $m_t$  experimentally. This is still very useful because the measured top mass from CDF [2],  $m_t = 174 \pm 10_{-12}^{+13}$  GeV, has a large error bars and D0 gives just the lower bound on  $m_t$ ,  $m_t \gtrsim 131$  GeV [3]. With the improved measurement for the  $Z$  partial width to  $b\bar{b}$ , primarily due to the use of new life time-based techniques, one may be able to put more precise bound on  $m_t$ . The experimental value for  $\Gamma(Z \rightarrow b\bar{b})$  has increased over the year, resulting in larger experimental value for  $R_b (\equiv \frac{\Gamma(Z \rightarrow b\bar{b})}{\Gamma(Z \rightarrow \text{hadrons})})$ , and therefore rather small upper bound on  $m_t$  is favored in the SM [1]. One could certainly interpret this as a possible manifestation of new physics beyond the SM, where at one loop the negative standard top quark contributions are cancelled to a certain extent by the contributions from the new particles, thereby allowing considerably larger  $m_t$  than in the SM. In fact, the minimal supersymmetric standard model (MSSM) realizes this possibility.

Another very interesting observable which encodes the one loop corrections to the  $Z \rightarrow b\bar{b}$  vertex is  $\epsilon_b$  first introduced in Ref. [4]. In supergravity(SUGRA) models, radiative electroweak symmetry breaking mechanism [5] can be described by at most 5 parameters: the top-quark mass ( $m_t$ ), the ratio of Higgs vacuum expectation values ( $\tan\beta$ ), and three universal soft-supersymmetry-breaking parameters ( $m_{1/2}, m_0, A$ )<sup>1</sup>.

In this work we explore the minimal  $SU(5)$  SUGRA [7] and the no-scale  $SU(5) \times U(1)$  SUGRA [8] in terms of  $\epsilon_b$  parameter which encodes the one-loop corrections to the  $Z \rightarrow b\bar{b}$  vertex. Moreover, we attempt to see how well these models can fit in rather uncomfortably

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<sup>1</sup>See, however, Ref. [6] for non-universal soft-supersymmetry breaking parameters

high 1993 LEP value for  $R_b$  ( $\equiv \frac{\Gamma(Z \rightarrow b\bar{b})}{\Gamma(Z \rightarrow \text{hadrons})}$ ).

## II. THE MINIMAL $SU(5)$ AND $SU(5) \times U(1)$ SUGRA MODELS

The minimal  $SU(5)$  and  $SU(5) \times U(1)$  SUGRA models both contain, at low energy, the SM gauge symmetry and the particle content of the MSSM. A few crucial differences between the two models are:

- (i) The unification groups are different,  $SU(5)$  versus  $SU(5) \times U(1)$ .
- (ii) The gauge coupling unification occurs at  $\sim 10^{16}$  GeV in the minimal  $SU(5)$  model whereas in  $SU(5) \times U(1)$  model it occurs at the string scale  $\sim 10^{18}$  GeV [9]. In  $SU(5) \times U(1)$  SUGRA, the gauge unification is delayed because of the effects of an additional pair of  $\mathbf{10}, \overline{\mathbf{10}}$  vector-like representations with intermediate-scale masses. The different heavy field content at the unification scale leads to different constraints from proton decay.
- (iii) In the minimal  $SU(5)$  SUGRA, proton decay is highly constraining whereas it is not in  $SU(5) \times U(1)$  SUGRA.

The procedure to restrict 5-dimensional parameter spaces is as follows [10]. First, upon sampling a specific choice of  $(m_{1/2}, m_0, A)$  at the unification scale and  $(m_t, \tan \beta)$  at the electroweak scale, the renormalization group equations (RGE) are run from the unification scale to the electroweak scale, where the radiative electroweak breaking condition is imposed by minimizing the effective 1-loop Higgs potential, which determines the Higgs mixing term  $\mu$  up to its sign. We also impose consistency constraints such as perturbative unification and the naturalness bound of  $m_{\tilde{g}} \lesssim 1$  TeV. Finally, all the known experimental bounds on the sparticle masses are imposed <sup>2</sup>. This procedure yields the restricted parameter spaces for the two models.

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<sup>2</sup>We use the following experimental lower bounds on the sparticle masses in GeV in the order of gluino, squarks, lighter stop, sleptons, and lighter chargino:  $m_{\tilde{g}} \gtrsim 150$ ,  $m_{\tilde{q}} \gtrsim 100$ ,  $m_{\tilde{t}_1} \gtrsim 45$ ,  $m_{\tilde{l}} \gtrsim 43$ ,  $m_{\chi_{\pm 1}^\pm} \gtrsim 45$ .

Further reduction in the number of input parameters in  $SU(5) \times U(1)$  SUGRA is made possible because in specific string-inspired scenarios for  $(m_{1/2}, m_0, A)$  at the unification scale these three parameters are computed in terms of just one of them [11]. One obtains  $m_0 = A = 0$  in the *no-scale* scenario and  $m_0 = \frac{1}{\sqrt{3}}m_{1/2}$ ,  $A = -m_{1/2}$  in the *dilaton* scenario<sup>3</sup>.

The low energy predictions for the sparticle mass spectra are quite different in the two SUGRA models mainly due to the different pattern of supersymmetry radiative breaking. In the minimal  $SU(5)$  SUGRA, all the squarks except the lighter stop and all the Higgs except the lighter neutral Higgs are quite heavy ( $\gtrsim$  a few hundred GeV) whereas they can be quite light in the  $SU(5) \times U(1)$  SUGRA. This difference leads to strikingly different phenomenology in the two models, for example in the flavor changing radiative decay  $b \rightarrow s\gamma$  [13].

### III. ONE-LOOP ELECTROWEAK RADIATIVE CORRECTIONS AND THE NEW $\epsilon$ PARAMETERS

There are several schemes to parametrize the electroweak vacuum polarization corrections [14–17]. It can be shown, by expanding the vacuum polarization tensors to order  $q^2$ , that one obtains three independent physical parameters. Alternatively, one can show that upon symmetry breaking three additional terms appear in the effective lagrangian [16]. In the  $(S, T, U)$  scheme [15], the deviations of the model predictions from the SM predictions (with fixed SM values for  $m_t, m_{H_{SM}}$ ) are considered as the effects from “new physics”. This scheme is only valid to the lowest order in  $q^2$ , and is therefore not applicable to a theory with light new particles comparable to  $M_Z$ . In the  $\epsilon$ -scheme [4,18], on the other hand, the model predictions are absolute and also valid up to higher orders in  $q^2$ , and therefore this scheme is more applicable to the electroweak precision tests of the MSSM [19] and a class

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<sup>3</sup>Note, however, that one loop correction changes this relation significantly [12].

of supergravity models [20].

There are two different  $\epsilon$ -schemes. The original scheme [18] was considered in one of author's previous analyses [20,21], where  $\epsilon_{1,2,3}$  are defined from a basic set of observables  $\Gamma_l, A_{FB}^l$  and  $M_W/M_Z$ . Due to the large  $m_t$ -dependent vertex corrections to  $\Gamma_b$ , the  $\epsilon_{1,2,3}$  parameters and  $\Gamma_b$  can be correlated only for a fixed value of  $m_t$ . Therefore,  $\Gamma_{tot}$ ,  $\Gamma_{hadron}$  and  $\Gamma_b$  were not included in Ref. [18]. However, in the new  $\epsilon$ -scheme, introduced recently in Ref. [4], the above difficulties are overcome by introducing a new parameter  $\epsilon_b$  to encode the  $Z \rightarrow b\bar{b}$  vertex corrections. The four  $\epsilon$ 's are now defined from an enlarged set of  $\Gamma_l$ ,  $\Gamma_b$ ,  $A_{FB}^l$  and  $M_W/M_Z$  without even specifying  $m_t$ . This new scheme was adopted in a previous analysis by one of us (G.P.) in the context of the  $SU(5) \times U(1)$  SUGRA models [22]. In this work we use this new  $\epsilon$ -scheme. As is well known, the SM contribution to  $\epsilon_1$  depends quadratically on  $m_t$  but only logarithmically on the SM Higgs boson mass ( $m_H$ ). Therefore upper bounds on  $m_t$  have a non-negligible  $m_H$  dependence: up to 20 GeV stronger when going from a heavy ( $\approx 1$  TeV) to a light ( $\approx 100$  GeV) Higgs boson. It is also known in the MSSM that the largest supersymmetric contributions to  $\epsilon_1$  are expected to arise from the  $\tilde{t}$ - $\tilde{b}$  sector, and in the limiting case of a very light stop, the contribution is comparable to that of the  $t$ - $b$  sector. The remaining squark, slepton, chargino, neutralino, and Higgs sectors all typically contribute considerably less. For increasing sparticle masses, the heavy sector of the theory decouples, and only SM effects with a *light* Higgs boson survive. However, for a light chargino ( $m_{\chi_1^\pm} \rightarrow \frac{1}{2}M_Z$ ), a  $Z$ -wavefunction renormalization threshold effect coming from  $Z$ -vacuum polarization diagram with the lighter chargino in the loop can introduce a substantial  $q^2$ -dependence in the calculation [19]. This results in a weaker upper bound on  $m_t$  than in the SM. The complete vacuum polarization contributions from the Higgs sector, the supersymmetric chargino-neutralino and sfermion sectors, and also the corresponding contributions in the SM have been included in our calculations [20]. However, the supersymmetric contributions to the non-oblique corrections except in  $\epsilon_b$  have been neglected.

Following Ref. [4],  $\epsilon_b$  is defined from  $\Gamma_b$ , the inclusive partial width for  $Z \rightarrow b\bar{b}$ , as

$$\epsilon_b = \frac{g_A^b}{g_A^l} - 1 \quad (1)$$

where  $g_A^b$  ( $g_A^l$ ) is the axial-vector coupling of  $Z$  to  $b$  ( $l$ ). In the SM, the diagrams for  $\epsilon_b$  involve top quarks and  $W^\pm$  bosons [23], and the contribution to  $\epsilon_b$  depends quadratically on  $m_t$  ( $\epsilon_b = -G_F m_t^2 / 4\sqrt{2}\pi^2 + \dots$ ). In supersymmetric models there are additional diagrams involving Higgs bosons and supersymmetric particles. The charged Higgs contributions have been calculated in Refs. [24,25] in the context of a non-supersymmetric two Higgs doublet model, and the contributions involving supersymmetric particles in Refs. [26,27]. The main features of the additional supersymmetric contributions are: (i) a negative contribution from charged Higgs–top exchange which grows as  $m_t^2 / \tan^2 \beta$  for  $\tan \beta \ll \frac{m_t}{m_b}$ ; (ii) a positive contribution from chargino–stop exchange which in this case grows as  $m_t^2 / \sin^2 \beta$ ; and (iii) a contribution from neutralino(neutral Higgs)–bottom exchange which grows as  $m_b^2 \tan^2 \beta$  and is negligible except for large values of  $\tan \beta$  (*i.e.*,  $\tan \beta \gtrsim \frac{m_t}{m_b}$ ) (the contribution (iii) has been neglected in our analysis).

#### IV. RESULTS AND DISCUSSION

In Figure 1 we present our numerical results for  $\epsilon_b$  in the two SUGRA models.  $\alpha_S(M_Z) = 0.118$  and  $m_b = 4.8$  GeV are used throughout the numerical calculations. We use the experimental value for  $\epsilon_b$ ,  $\epsilon_b^{exp} = (0.9 \pm 4.2) \times 10^{-3}$ , determined from the latest  $\epsilon$ -analysis using the LEP and SLC data in Ref. [28]. The discontinuity in the chargino mass in the minimal model in the figure is simply due to the use of large steps in sampling the value of  $m_{1/2}$ . The values of  $m_t$  are chosen in such a way that the approximate  $\epsilon_b$ -deduced  $m_t$  bounds are readily obtained from the figure. Only one value of  $m_t$  is displayed in the  $SU(5) \times U(1)$  SUGRA because considerable portion of the model predictions are overlapped for two different values of  $m_t$  due to the steep rise in  $\epsilon_b$  for a light chargino. The reason why the rise in  $\epsilon_b$  in the  $SU(5) \times U(1)$  SUGRA is much steeper than in the minimal  $SU(5)$  SUGRA is that the stop mass scales with the chargino mass in the no-scale model whereas

it does not in the minimal model. Therefore, the light chargino effect in  $\epsilon_b$  is optimized better in the no-scale  $SU(5) \times U(1)$  SUGRA. This difference leads to different  $\epsilon_b$ -deduced  $m_t$  bounds in the two models. The approximate bounds at 90% C. L. are  $m_t \lesssim 175$  (185) GeV for the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. In the no-scale model, one can also determine the sign on  $\mu$  to be positive for  $m_t \gtrsim 180$  GeV. The lowest value of  $\epsilon_b$  for a fixed  $m_t$  represents the lowest  $\tan \beta$  for not too large  $\tan \beta$ <sup>4</sup>. It is  $\tan \beta = 1.5$  (4.0) for the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA in Figure 1. From this, we obtain low  $\tan \beta - m_t$  correlated bounds at 90% C. L., which are for  $\tan \beta \lesssim 2.5$  (4.0),  $m_t \lesssim 170$  (180) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. Although the  $m_t$  values from CDF have rather large error bars at present, one can imagine an interesting situation in the near future where the  $m_t$  values from CDF turns out to fall between the above  $\epsilon_b$ -deduced  $m_t$  bounds, disfavoring only one model.

The experimental value for  $R_b(\equiv \frac{\Gamma(Z \rightarrow b\bar{b})}{\Gamma(Z \rightarrow \text{hadrons})})$  from the 1993 LEP data are reported very recently to be rather high,  $0.2192 \pm 0.0018$ , in comparison with the SM predictions [1]. In an attempt to see how much the situation can improve in SUGRA models, we now calculate  $R_b$  in the two SUGRA models<sup>5</sup>. In Figure 2 we show the model predictions for  $R_b$  in the two models. As seen in the figure, the  $R_b$  constraint is much stronger than the  $\epsilon_b$  constraint. The  $R_b$ -deduced  $m_t$  bounds at 90% C. L. are  $m_t \lesssim 165$  (175) GeV for the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. From the figure, one can also put bounds on the chargino mass, which are  $m_{\chi_1^\pm} \lesssim 85$  (70) GeV for  $m_t \gtrsim 160$  (170) GeV for the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. Similarly, one can also obtain bounds on the lighter stop mass given by  $m_{\tilde{t}_1} \lesssim 500$  (190) GeV for  $m_t \gtrsim 160$  (170) GeV for the minimal  $SU(5)$  (no-scale

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<sup>4</sup>For large  $\tan \beta (\gtrsim \frac{m_t}{m_b})$ , the charged Higgs diagram gets a significant contribution proportional to  $-m_b^2 \tan^2 \beta$  coming from the charged Higgs coupling to  $b_R$ , thereby driving  $\epsilon_b$  even below the value corresponding to the lowest  $\tan \beta$ .

<sup>5</sup>We use the expression for  $R_b$  in terms of  $\epsilon$ 's given in Ref. [4]



$SU(5) \times U(1)$  SUGRA. Therefore, in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA, if the top turns out to be heavier than 160 (170) GeV, then only the lighter chargino may be detected at LEP II. The  $\tan\beta$ -dependence is very pronounced in the no-scale model for  $\mu > 0$ . The low values of  $\tan\beta$  are as indicated in the figure. For  $\tan\beta = 2$ , the dotted curve becomes nearly flat as the chargino mass becomes large. This is because the charged Higgs contribution nearly cancels the chargino contribution [26], making  $R_b$  get saturated much faster to the SM value. As in the  $\epsilon_b$  constraint above, the low  $\tan\beta - m_t$  correlated bounds at 90% C. L. are obtained as follows: for  $\tan\beta \lesssim 4.0$ ,  $m_t \lesssim 160$  (170) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. In the no-scale model,  $\tan\beta \lesssim 2$  is excluded even at 95% C. L. for  $m_t \gtrsim 170$  GeV. From  $R_b$ , one can also determine  $\mu$  to be positive in both models. It is very interesting for one to see that the low- $\tan\beta$  region is severely constrained by both constraints above. We would like to stress here the fact that our calculations are fairly accurate in the low- $\tan\beta$  region because the diagrams neglected in the calculations can be safely neglected there. The major features of the constraints from  $\epsilon_b$  and  $R_b$  for the two SUGRA models are summarized in the Table 1.

## V. CONCLUSIONS

We have computed the one-loop electroweak corrections to the  $Z \rightarrow b\bar{b}$  vertex in terms of  $\epsilon_b$  and  $R_b$  in the context of the minimal  $SU(5)$  and no-scale  $SU(5) \times U(1)$  supergravity models. We use the latest LEP data on  $\epsilon_b$  and  $R_b$  in order to constrain further the two models. We find that the present LEP data on  $\epsilon_b$  and  $R_b$  constrain the two models rather severely. Especially, the low- $\tan\beta$  region is constrained more severely.  $\tan\beta \lesssim 2.5$  (4.0) is excluded by  $\epsilon_b$  at 90% C. L. for  $m_t \gtrsim 170$  (180) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. Even more stringent constraint comes from  $R_b$ . It excludes  $\tan\beta \lesssim 4.0$  at 90% C. L. for  $m_t \gtrsim 160$  (170) GeV in the minimal  $SU(5)$  (no-scale  $SU(5) \times U(1)$ ) SUGRA. We also find that the sign on  $\mu$  in the two models can be determined from  $\epsilon_b$  and  $R_b$  at 90% C. L. This can be of special interest in the minimal  $SU(5)$  because the low- $\tan\beta$  region is

phenomenologically favored by the measured ratio  $m_b/m_\tau$ . We also find that the sign on  $\mu$  in the two models can be determined from  $\epsilon_b$  and  $R_b$  at 90% C. L.

With improved measurement on the top mass by CDF in the near future, there may be an amusing possibility that one could favor one model over the other from the  $Z \rightarrow b\bar{b}$  constraints. And also, in the no-scale  $SU(5) \times U(1)$  SUGRA, if the top turns out to be heavier than 170 GeV, then only the lighter chargino lighter than 80 GeV may be detected at LEP II.

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## REFERENCES

- [1] D. Schaile, talk given at 27th International Conference on High Energy Physics, Glasgow, July 1994.
- [2] CDF Collaboration, Phys. Rev. Lett. **73** (1994) 225.
- [3] D0 Collaboration, Phys. Rev. Lett. **72** (1994) 2138.
- [4] G. Altarelli, R. Barbieri, and F. Caravaglios, Nucl. Phys. B **405** (1993) 3.
- [5] L. Ibáñez and G. Ross, Phys. Lett. B **110** (1982) 215; K. Inoue, *et al.*, Prog. Theor. Phys. 68 (1982) 927; L. Ibáñez, Nucl. Phys. B **218** (1983) 514 and Phys. Lett. B **118** (1982) 73; H. P. Nilles, Nucl. Phys. B **217** (1983) 366; J. Ellis, D. V. Nanopoulos, and K. Tamvakis, Phys. Lett. B **121** (1983) 123; J. Ellis, J. Hagelin, D. V. Nanopoulos, and K. Tamvakis, Phys. Lett. B **125** (1983) 275; L. Alvarez-Gaumé, J. Polchinski, and M. Wise, Nucl. Phys. B **221** (1983) 495; L. Ibáñez and C. López, Phys. Lett. B **126** (1983) 54 and Nucl. Phys. B **233** (1984) 545; C. Kounnas, A. Lahanas, D. V. Nanopoulos, and M. Quirós, Phys. Lett. B **132** (1983) 95 and C. Kounnas, A. Lahanas, D. V. Nanopoulos, and M. Quirós, Nucl. Phys. B **236** (1984) 438.
- [6] D. Matalliotakis and H. P. Nilles, TUM-HEP-201/94.
- [7] For reviews see R. Arnowitt and P. Nath, *Applied N=1 Supergravity* (World Scientific, Singapore 1983); H. P. Nilles, Phys. Rep. **110** (1984) 1.
- [8] For a recent review see J. L. Lopez, D. V. Nanopoulos, and A. Zichichi, CERN-TH.6926/93 (unpublished).
- [9] J. L. Lopez, D. V. Nanopoulos, and A. Zichichi, Phys. Rev. D **49** (1994) 343 and references therein.
- [10] S. Kelley, J. L. Lopez, D. V. Nanopoulos, H. Pois, and K. Yuan, Nucl. Phys. B **398** (1993) 3.

- [11] See *e.g.*, L. Ibáñez and D. Lüst, Nucl. Phys. B **382** (1992) 305; V. Kaplunovsky and J. Louis, Phys. Lett. B **306** (1993) 269; A. Brignole, L. Ibáñez, and C. Muñoz, FTUAM-26/93.
- [12] K. Choi, J. E. Kim and H. P. Nilles, SNUTP-94-19 (1994).
- [13] J. L. Lopez, D. V. Nanopoulos, and G. T. Park, Phys. Rev. D **48** (1993) R974.
- [14] D. Kennedy and B. Lynn, Nucl. Phys. B **322** (1989) 1; D. Kennedy, B. Lynn, C. Im, and R. Stuart, Nucl. Phys. B **321** (1989) 83.
- [15] M. Peskin and T. Takeuchi, Phys. Rev. Lett. **65** (1990) 964; W. Marciano and J. Rosner, Phys. Rev. Lett. **65** (1990) 2963; D. Kennedy and P. Langacker, Phys. Rev. Lett. **65** (1990) 2967.
- [16] B. Holdom and J. Terning, Phys. Lett. B **247** (1990) 88; M. Golden and L. Randall, Nucl. Phys. B **361** (1991) 3; A. Dobado, D. Espriu, and M. Herrero, Phys. Lett. B **255** (1991) 405.
- [17] G. Altarelli and R. Barbieri, Phys. Lett. B **253** (1990) 161
- [18] G. Altarelli, R. Barbieri, and S. Jadach, Nucl. Phys. B **369** (1992) 3.
- [19] R. Barbieri, M. Frigeni, and F. Caravaglios, Phys. Lett. B **279** (1992) 169.
- [20] J. L. Lopez, D. V. Nanopoulos, G. T. Park, H. Pois, and K. Yuan, Phys. Rev. D **48** (1993) 3297.
- [21] J. L. Lopez, D. V. Nanopoulos, G. T. Park, and A. Zichichi, Phys. Rev. D **49** (1994) 355.
- [22] J. L. Lopez, D. V. Nanopoulos, G. T. Park, and A. Zichichi, Phys. Rev. D **49** (1994) 4835.
- [23] J. Bernabeu, A. Pich, and A. Santamaria, Phys. Lett. B **200** (1988) 569; W. Beenaker

- and W. Hollik, Z. Phys. C40, 141(1988); A. Akhundov, D. Bardin, and T. Riemann, Nucl. Phys. B **276** (1986) 1; F. Boudjema, A. Djouadi, and C. Verzegnassi, Phys. Lett. B **238** (1990) 423.
- [24] A. Denner, R. Guth, W. Hollik, and J. Kühn, Z. Phys. C51, 695(1991). The neutral Higgs contributions to  $Z \rightarrow b\bar{b}$  were also calculated here.
- [25] G. T. Park, Mod. Phys. Lett. A **9** (1994) 321.
- [26] M. Boulware and D. Finnell, Phys. Rev. D **44** (1991) 2054.
- [27] A. Djouadi, G. Girardi, C. Verzegnassi, W. Hollik, and F. Renard, Nucl. Phys. B **349** (1991) 48.
- [28] R. Barbieri, talk given at Rencontres Physique de la Valle d'Aosta, La Thuile, IFUP-TH 28/94 (1994).

## FIGURES

FIG. 1. The predictions for  $\epsilon_b$  versus the lighter chargino mass in the minimal  $SU(5)$  SUGRA for  $m_t = 160, 175$  GeV (top row) and in the no-scale  $SU(5) \times U(1)$  SUGRA for  $m_t = 180$  GeV (bottom row). The values of  $m_t$  are as indicated. The points above the horizontal solid line are allowed at 90% C.L.

FIG. 2. The predictions for  $R_b$  versus the lighter chargino mass in the minimal  $SU(5)$  SUGRA for  $m_t = 160$  GeV (top row) and in the no-scale  $SU(5) \times U(1)$  SUGRA for  $m_t = 170$  GeV (bottom row). The values of  $\tan \beta$  are as indicated near the dotted curves (bottom row). The points above the horizontal solid lines are allowed at 90 or 95% C.L.

# TABLES

TABLE I. The major features of the constraints from  $\epsilon_b$  and  $R_b$  for the two SUGRA models considered.

	Minimal $SU(5)$	no-scale $SU(5) \times U(1)$
$\epsilon_b$ (90% C.L.)	$m_t \lesssim 175$ GeV for any $\tan \beta$	$m_t \lesssim 185$ GeV for any $\tan \beta$
	$m_t \lesssim 170$ GeV for $\tan \beta \lesssim 2.5$	$m_t \lesssim 180$ GeV for $\tan \beta \lesssim 4$
$R_b$ (90% C.L.)	$m_t \lesssim 165$ GeV for any $\tan \beta$	$m_t \lesssim 175$ GeV for any $\tan \beta$
	$m_t \lesssim 160$ GeV for $\tan \beta \lesssim 4$	$m_t \lesssim 170$ GeV for $\tan \beta \lesssim 4$
	For $m_t \gtrsim 160$ GeV,	For $m_t \gtrsim 170$ GeV,
	$m_{\chi_1^\pm} \lesssim 85$ GeV and $m_{\tilde{t}_1} \lesssim 500$ GeV	$m_{\chi_1^\pm} \lesssim 70$ GeV and $m_{\tilde{t}_1} \lesssim 190$ GeV

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This figure "fig1-2.png" is available in "png" format from:

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